

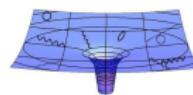
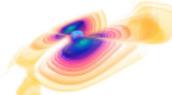
Spin effects in the post-Newtonian approximation: Quadrupole deformation of Neutron stars and three-body interactions

J. Hartung S. Hergt G. Schäfer J. Steinhoff

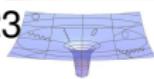


Theoretisch-Physikalisches Institut
Friedrich-Schiller-Universität Jena

Video seminar of the SFB/TR7, November 22nd, 2010



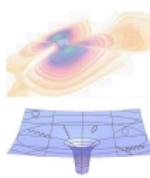
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DFG : SFB/TR7 “Gravitational Wave Astronomy” and GRK 1523

Outline

- ① Three-body interactions with spin:
NLO Spin-Orbit and NLO Spin(a)-Spin(b) Hamiltonians
 - 📄 J. Hartung and J. Steinhoff
Phys. Rev. D, submitted, arXiv:1011.1179
- ② Action approach to canonical formulation of spin in GR
 - 📄 J. Steinhoff and G. Schäfer
Europhys. Lett. **87**, 50004 (2009)
- ③ Quadrupole deformation of Neutron stars due to Spin:
NLO Spin(1)-Spin(1) Hamiltonian
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Class. Quant. Grav. **27**, 135007 (2010)



Three-body interactions

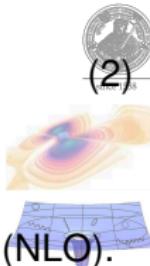
- $H^{\text{ADM}} \doteq \text{ADM energy expressed in terms of canonical variables}$ after field constraints are solved in the ADMTT gauge.
- Canonical variables are denoted by a hat $\hat{\cdot}$.
- Only pairwise two-body interactions in the Newtonian case:

$$H^N = \sum_a \frac{\hat{\mathbf{p}}_a^2}{2m_a} - \frac{1}{2} \sum_a \sum_{b \neq a} \frac{Gm_a m_b}{\hat{r}_{ab}} \quad (1)$$

- At 1PN three-body interactions appear:

$$\begin{aligned} H^{1\text{PN}} = & - \sum_a \frac{(\hat{\mathbf{p}}_a^2)^2}{8m_a^3} + \sum_a \sum_{b \neq a} \frac{G}{\hat{r}_{ab}} \left[-\frac{3m_b}{2m_a} \hat{\mathbf{p}}_a^2 + \frac{1}{4} (7(\hat{\mathbf{p}}_a \hat{\mathbf{p}}_b) + (\hat{\mathbf{n}}_{ab} \hat{\mathbf{p}}_a)(\hat{\mathbf{n}}_{ab} \hat{\mathbf{p}}_b)) \right] \\ & + \sum_a \sum_{b \neq a} \frac{G^2 m_a^2 m_b}{2\hat{r}_{ab}^2} + \sum_a \sum_{b \neq a} \sum_{c \neq a,b} \frac{G^2 m_a m_b m_c}{2\hat{r}_{ab} \hat{r}_{ac}} \end{aligned} \quad (2)$$

- The leading order in spin is only a sum of two-body interactions.
- Three-body interactions with spin appear at next-to-leading order (NLO).



NLO Spin-Orbit, two-body interaction part

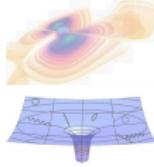
Hamiltonian first derived: Damour, Jaranowski, Schäfer (2008)

See also: Tagoshi, Ohashi, Owen (2001); Faye, Blanchet, Buonanno (2006)

$$H_{\text{SO},[2]}^{\text{NLO}} = \sum_a \sum_{b \neq a} \left(-G \frac{((\hat{\mathbf{p}}_a \times \hat{\mathbf{S}}_a) \cdot \hat{\mathbf{n}}_{ab})}{\hat{r}_{ab}^2} \left[\frac{5m_b \hat{\mathbf{p}}_a^2}{8m_a^3} + \frac{3(\hat{\mathbf{p}}_a \cdot \hat{\mathbf{p}}_b)}{4m_a^2} - \frac{3\hat{\mathbf{p}}_b^2}{4m_a m_b} \right. \right. \right. \\ \left. \left. \left. + \frac{3(\hat{\mathbf{p}}_a \cdot \hat{\mathbf{n}}_{ab})(\hat{\mathbf{p}}_b \cdot \hat{\mathbf{n}}_{ab})}{4m_a^2} + \frac{3(\hat{\mathbf{p}}_b \cdot \hat{\mathbf{n}}_{ab})^2}{2m_a m_b} \right] \right. \right. \\ \left. \left. + G \frac{((\hat{\mathbf{p}}_b \times \hat{\mathbf{S}}_a) \cdot \hat{\mathbf{n}}_{ab})}{\hat{r}_{ab}^2} \left[\frac{(\hat{\mathbf{p}}_a \cdot \hat{\mathbf{p}}_b)}{m_a m_b} + \frac{3(\hat{\mathbf{p}}_a \cdot \hat{\mathbf{n}}_{ab})(\hat{\mathbf{p}}_b \cdot \hat{\mathbf{n}}_{ab})}{m_a m_b} \right] \right. \right. \\ \left. \left. + G \frac{((\hat{\mathbf{p}}_a \times \hat{\mathbf{S}}_a) \cdot \hat{\mathbf{p}}_b)}{\hat{r}_{ab}^2} \left[\frac{2(\hat{\mathbf{p}}_b \cdot \hat{\mathbf{n}}_{ab})}{m_a m_b} - \frac{3(\hat{\mathbf{p}}_a \cdot \hat{\mathbf{n}}_{ab})}{4m_a^2} \right] \right. \right. \\ \left. \left. - G^2 \frac{((\hat{\mathbf{p}}_a \times \hat{\mathbf{S}}_a) \cdot \hat{\mathbf{n}}_{ab})}{\hat{r}_{ab}^3} \left[\frac{11m_b}{2} + \frac{5m_b^2}{m_a} \right] \right. \right. \\ \left. \left. + G^2 \frac{((\hat{\mathbf{p}}_b \times \hat{\mathbf{S}}_a) \cdot \hat{\mathbf{n}}_{ab})}{\hat{r}_{ab}^3} \left[6m_a + \frac{15m_b}{2} \right] \right) \right)$$



since 1558

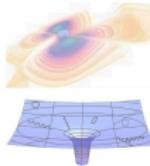
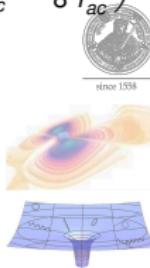


NLO Spin-Orbit, three-body interaction part

Hartung, Steinhoff, arXiv:1011.1179

$$\begin{aligned}
 H_{SO,[3]}^{\text{NLO}} = & \sum_a \sum_{b \neq a} \sum_{c \neq a,b} \left(G^2 m_a \left[\frac{1}{\hat{s}_{abc}} \left\{ \left(\frac{3}{\hat{r}_{ab}\hat{r}_{ac}} - \frac{6}{\hat{r}_{ab}\hat{r}_{bc}} - \frac{3}{\hat{r}_{ac}\hat{r}_{bc}} - \frac{3\hat{r}_{ab}}{\hat{r}_{ac}^2\hat{r}_{bc}} + \frac{3\hat{r}_{bc}}{\hat{r}_{ab}\hat{r}_{ac}^2} \right) ((\hat{\mathbf{n}}_{ac} \times \hat{\mathbf{p}}_b) \hat{\mathbf{S}}_c) \right. \right. \right. \right. \\
 & + \left(-\frac{8}{\hat{r}_{bc}^2} + \frac{8}{\hat{r}_{ab}\hat{r}_{ac}} - \frac{4}{\hat{r}_{ab}\hat{r}_{bc}} - \frac{4}{\hat{r}_{ac}\hat{r}_{bc}} - \frac{4\hat{r}_{ab}}{\hat{r}_{ac}\hat{r}_{bc}^2} - \frac{4\hat{r}_{ac}}{\hat{r}_{ab}\hat{r}_{bc}^2} \right) ((\hat{\mathbf{n}}_{bc} \times \hat{\mathbf{p}}_b) \hat{\mathbf{S}}_c) \Big\} \\
 & + \frac{((\hat{\mathbf{n}}_{ac} \times \hat{\mathbf{n}}_{bc}) \hat{\mathbf{S}}_c)}{\hat{s}_{abc}^2} \left\{ \left(\frac{4}{\hat{r}_{ab}} - \frac{2\hat{r}_{bc}}{\hat{r}_{ab}\hat{r}_{ac}} + \frac{2\hat{r}_{ac}}{\hat{r}_{ab}^2} - \frac{2\hat{r}_{bc}^2}{\hat{r}_{ab}^2\hat{r}_{ac}} + \frac{8\hat{r}_{ab}}{\hat{r}_{ac}^2} + \frac{7\hat{r}_{bc}}{\hat{r}_{ac}^2} - \frac{2\hat{r}_{bc}^2}{\hat{r}_{ab}\hat{r}_{ac}^2} + \frac{\hat{r}_{bc}}{\hat{r}_{ab}^2} - \frac{\hat{r}_{bc}^3}{\hat{r}_{ab}^2\hat{r}_{ac}^2} \right) (\hat{\mathbf{n}}_{ab} \hat{\mathbf{p}}_b) \right. \\
 & + \left. \left. \left. \left. + \left(\frac{12}{\hat{r}_{ab}} - \frac{2}{\hat{r}_{ac}} + \frac{5}{\hat{r}_{bc}} - \frac{2\hat{r}_{ab}}{\hat{r}_{ac}^2} + \frac{7\hat{r}_{bc}}{\hat{r}_{ac}^2} - \frac{2\hat{r}_{ab}}{\hat{r}_{ac}\hat{r}_{bc}} + \frac{6\hat{r}_{ac}}{\hat{r}_{ab}\hat{r}_{bc}} - \frac{\hat{r}_{ab}^2}{\hat{r}_{ac}^2\hat{r}_{bc}} + \frac{8\hat{r}_{bc}^2}{\hat{r}_{ab}\hat{r}_{ac}^2} \right) (\hat{\mathbf{n}}_{bc} \hat{\mathbf{p}}_b) + \frac{16(\hat{\mathbf{n}}_{ac} \hat{\mathbf{p}}_b)}{\hat{r}_{ab}} \right) \right] \right. \\
 & + G^2 \frac{m_a m_b}{m_c} \left[\frac{1}{\hat{s}_{abc}^2} \left\{ \frac{1}{\hat{r}_{ab}^3} \left(-2\hat{r}_{ac}^2 + \hat{r}_{bc}^2 - \frac{3}{2}\hat{r}_{ac}\hat{r}_{bc} + \frac{1}{2}\frac{\hat{r}_{bc}^3}{\hat{r}_{ac}} \right) + \frac{1}{\hat{r}_{ab}^2} \left(-4\hat{r}_{ac} + \hat{r}_{bc} + \frac{\hat{r}_{bc}^2}{\hat{r}_{ac}} \right) - \frac{1}{\hat{r}_{ac}} \right. \right. \\
 & - \frac{1}{\hat{r}_{ab}} \left(2 + \frac{1}{2}\frac{\hat{r}_{bc}}{\hat{r}_{ac}} \right) \Big\} (\hat{\mathbf{n}}_{ac} \hat{\mathbf{p}}_c) ((\hat{\mathbf{n}}_{ac} \times \hat{\mathbf{n}}_{bc}) \hat{\mathbf{S}}_c) + \frac{1}{\hat{s}_{abc}} \left\{ \frac{1}{\hat{r}_{ab}^3} \left(\frac{1}{8}\hat{r}_{ac} - \frac{5}{8}\hat{r}_{bc} + \frac{3}{4}\frac{\hat{r}_{ac}^2}{\hat{r}_{bc}} - \frac{1}{8}\frac{\hat{r}_{bc}^2}{\hat{r}_{ac}} - \frac{1}{8}\frac{\hat{r}_{bc}^3}{\hat{r}_{ac}^2} \right) \right. \\
 & + \frac{1}{\hat{r}_{ab}^2} \left(-\frac{5}{8} + \frac{3}{4}\frac{\hat{r}_{ac}}{\hat{r}_{bc}} - \frac{1}{8}\frac{\hat{r}_{bc}^2}{\hat{r}_{ac}^2} \right) + \frac{1}{\hat{r}_{ab}} \left(\frac{3}{8\hat{r}_{ac}} - \frac{1}{\hat{r}_{bc}} + \frac{3}{8}\frac{\hat{r}_{bc}}{\hat{r}_{ac}^2} \right) - \frac{\hat{r}_{ab}}{4\hat{r}_{ac}^2\hat{r}_{bc}} - \frac{1}{4\hat{r}_{ac}\hat{r}_{bc}} \\
 & \left. \left. + \frac{1}{8\hat{r}_{ac}^2} \right\} ((\hat{\mathbf{n}}_{ac} \times \hat{\mathbf{p}}_c) \hat{\mathbf{S}}_c) + (a \leftrightarrow b) \right] - \frac{G^2}{\hat{r}_{ab}^2} \left(\frac{5}{\hat{r}_{ac}} + \frac{1}{\hat{r}_{bc}} \right) \frac{m_b m_c}{m_a} ((\hat{\mathbf{n}}_{ab} \times \hat{\mathbf{p}}_a) \hat{\mathbf{S}}_a) \Big)
 \end{aligned}$$

$$\hat{\mathbf{S}}_{abc} = \hat{r}_{ab} + \hat{r}_{ac} + \hat{r}_{bc}$$



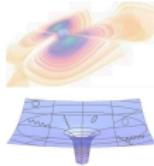
NLO Spin(a)-Spin(b), two-body interaction part

Partial result: Porto, Rothstein (2006). Full result: Steinhoff, Hergt, Schäfer (2008).

$$\begin{aligned} H_{S_a S_b, [2]}^{\text{NLO}} = & \sum_a \sum_{b \neq a} \left(\frac{G}{4m_a m_b \hat{r}_{ab}^3} \left[\frac{3}{2} ((\hat{\mathbf{p}}_a \times \hat{\mathbf{S}}_a) \cdot \hat{\mathbf{n}}_{ab}) ((\hat{\mathbf{p}}_b \times \hat{\mathbf{S}}_b) \cdot \hat{\mathbf{n}}_{ab}) \right. \right. \\ & + 6 ((\hat{\mathbf{p}}_b \times \hat{\mathbf{S}}_a) \cdot \hat{\mathbf{n}}_{ab}) ((\hat{\mathbf{p}}_a \times \hat{\mathbf{S}}_b) \cdot \hat{\mathbf{n}}_{ab}) - \frac{1}{2} (\hat{\mathbf{S}}_a \cdot \hat{\mathbf{p}}_b) (\hat{\mathbf{S}}_b \cdot \hat{\mathbf{p}}_a) \\ & - 15 (\hat{\mathbf{S}}_a \cdot \hat{\mathbf{n}}_{ab}) (\hat{\mathbf{S}}_b \cdot \hat{\mathbf{n}}_{ab}) (\hat{\mathbf{p}}_a \cdot \hat{\mathbf{n}}_{ab}) (\hat{\mathbf{p}}_b \cdot \hat{\mathbf{n}}_{ab}) + (\hat{\mathbf{S}}_a \cdot \hat{\mathbf{p}}_a) (\hat{\mathbf{S}}_b \cdot \hat{\mathbf{p}}_b) \\ & - 3 (\hat{\mathbf{S}}_a \cdot \hat{\mathbf{n}}_{ab}) (\hat{\mathbf{S}}_b \cdot \hat{\mathbf{n}}_{ab}) (\hat{\mathbf{p}}_a \cdot \hat{\mathbf{p}}_b) + 3 (\hat{\mathbf{S}}_a \cdot \hat{\mathbf{p}}_b) (\hat{\mathbf{S}}_b \cdot \hat{\mathbf{n}}_{ab}) (\hat{\mathbf{p}}_a \cdot \hat{\mathbf{n}}_{ab}) \\ & + 3 (\hat{\mathbf{S}}_b \cdot \hat{\mathbf{p}}_a) (\hat{\mathbf{S}}_a \cdot \hat{\mathbf{n}}_{ab}) (\hat{\mathbf{p}}_b \cdot \hat{\mathbf{n}}_{ab}) + 3 (\hat{\mathbf{S}}_a \cdot \hat{\mathbf{p}}_a) (\hat{\mathbf{S}}_b \cdot \hat{\mathbf{n}}_{ab}) (\hat{\mathbf{p}}_b \cdot \hat{\mathbf{n}}_{ab}) \\ & + 3 (\hat{\mathbf{S}}_b \cdot \hat{\mathbf{p}}_b) (\hat{\mathbf{S}}_a \cdot \hat{\mathbf{n}}_{ab}) (\hat{\mathbf{p}}_a \cdot \hat{\mathbf{n}}_{ab}) - 3 (\hat{\mathbf{S}}_a \cdot \hat{\mathbf{S}}_b) (\hat{\mathbf{p}}_a \cdot \hat{\mathbf{n}}_{ab}) (\hat{\mathbf{p}}_b \cdot \hat{\mathbf{n}}_{ab}) \\ & \left. \left. + \frac{1}{2} (\hat{\mathbf{S}}_a \cdot \hat{\mathbf{S}}_b) (\hat{\mathbf{p}}_a \cdot \hat{\mathbf{p}}_b) \right] \right. \\ & + \frac{3G}{2m_a^2 \hat{r}_{ab}^3} \left[- ((\hat{\mathbf{p}}_a \times \hat{\mathbf{S}}_a) \cdot \hat{\mathbf{n}}_{ab}) ((\hat{\mathbf{p}}_a \times \hat{\mathbf{S}}_b) \cdot \hat{\mathbf{n}}_{ab}) \right. \\ & \quad \left. + (\hat{\mathbf{S}}_a \cdot \hat{\mathbf{S}}_b) (\hat{\mathbf{p}}_a \cdot \hat{\mathbf{n}}_{ab})^2 - (\hat{\mathbf{S}}_a \cdot \hat{\mathbf{n}}_{ab}) (\hat{\mathbf{S}}_b \cdot \hat{\mathbf{p}}_a) (\hat{\mathbf{p}}_a \cdot \hat{\mathbf{n}}_{ab}) \right] \\ & \left. + \frac{6G^2 m_a}{\hat{r}_{ab}^4} [(\hat{\mathbf{S}}_a \cdot \hat{\mathbf{S}}_b) - 2 (\hat{\mathbf{S}}_a \cdot \hat{\mathbf{n}}_{ab}) (\hat{\mathbf{S}}_b \cdot \hat{\mathbf{n}}_{ab})] \right) \end{aligned}$$



since 1558



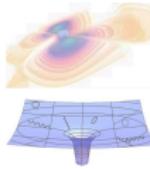
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Hartung, Steinhoff, arXiv:1011.1179

$$\begin{aligned}
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 & + \frac{2}{\hat{r}_{ab}^3} \left(2\hat{r}_{ac} + \frac{\hat{r}_{ac}^2}{\hat{r}_{bc}} \right) \left. \right\} + (\hat{\mathbf{n}}_{bc} \hat{\mathbf{S}}_a) (\hat{\mathbf{n}}_{ac} \hat{\mathbf{S}}_b) \left\{ -\frac{1}{\hat{r}_{ac}^2} - \frac{1}{\hat{r}_{ac} \hat{r}_{bc}} - \frac{2\hat{r}_{ab}}{\hat{r}_{ac}^2 \hat{r}_{bc}} - \frac{\hat{r}_{ab}^2}{2\hat{r}_{ac}^2 \hat{r}_{bc}^2} + \frac{\hat{r}_{ac}}{\hat{r}_{ab} \hat{r}_{bc}^2} \right. \\
 & + \frac{2}{\hat{r}_{ab}^2} \left(-1 + \frac{\hat{r}_{ac}}{\hat{r}_{bc}} + \frac{\hat{r}_{ac}^2}{\hat{r}_{bc}^2} \right) + \frac{1}{\hat{r}_{ab}^3} \left(2\hat{r}_{ac} + \frac{2\hat{r}_{ac}^2}{\hat{r}_{bc}} + \frac{\hat{r}_{ac}^3}{\hat{r}_{bc}^2} \right) + \frac{3(\hat{r}_{ac} + \hat{r}_{bc})^2}{\hat{r}_{ab}^4} + \frac{3(\hat{r}_{ac} + \hat{r}_{bc})^3}{2\hat{r}_{ab}^5} \left. \right\} \\
 & + (\hat{\mathbf{n}}_{ac} \hat{\mathbf{S}}_a) (\hat{\mathbf{n}}_{ac} \hat{\mathbf{S}}_b) \left\{ -\frac{3}{\hat{r}_{ab}^5} \left(3\hat{r}_{ac}^3 + 3\hat{r}_{ac}^2 \hat{r}_{bc} + \hat{r}_{ac} \hat{r}_{bc}^2 + \frac{\hat{r}_{ac}^4}{\hat{r}_{bc}} \right) - \frac{6}{\hat{r}_{ab}^4} \left(2\hat{r}_{ac}^2 + \hat{r}_{ac} \hat{r}_{bc} + \frac{\hat{r}_{ac}^3}{\hat{r}_{bc}} \right) \right. \\
 & - \frac{1}{\hat{r}_{ab}^3} \left(2\hat{r}_{ac} + 2\hat{r}_{bc} + \frac{\hat{r}_{ac}^2}{\hat{r}_{bc}} + \frac{\hat{r}_{bc}^2}{\hat{r}_{ac}} \right) - \frac{2}{\hat{r}_{ab}^2} \left(1 - \frac{2\hat{r}_{ac}}{\hat{r}_{bc}} + \frac{\hat{r}_{bc}}{\hat{r}_{ac}} \right) + \frac{1}{\hat{r}_{ab}} \left(\frac{1}{\hat{r}_{ac}} + \frac{4}{\hat{r}_{bc}} \right) + \frac{2}{\hat{r}_{ac} \hat{r}_{bc}} \left. \right\} \\
 & + (\hat{\mathbf{n}}_{ac} \hat{\mathbf{S}}_a) (\hat{\mathbf{n}}_{bc} \hat{\mathbf{S}}_b) \left\{ -\frac{2}{\hat{r}_{ab}^2} + \frac{\hat{r}_{ac}}{\hat{r}_{ab}^3} + \frac{3(\hat{r}_{ac} + \hat{r}_{bc})^2}{\hat{r}_{ab}^4} + \frac{3(\hat{r}_{ac} + \hat{r}_{bc})^3}{2\hat{r}_{ab}^5} \right\} + (\hat{\mathbf{S}}_a \hat{\mathbf{S}}_b) \left\{ \frac{2}{\hat{r}_{ac}^2} - \frac{3}{2\hat{r}_{ac} \hat{r}_{bc}} \right. \\
 & + \frac{3}{2} \frac{\hat{r}_{ac}}{\hat{r}_{bc}^3} + \hat{r}_{ab} \left(\frac{3}{2\hat{r}_{ac}^3} + \frac{1}{\hat{r}_{ac}^2 \hat{r}_{bc}} \right) - \frac{\hat{r}_{ab}^2}{2\hat{r}_{ac}^2 \hat{r}_{bc}^2} - \frac{\hat{r}_{ab}^3}{\hat{r}_{ac}^3 \hat{r}_{bc}^2} - \frac{\hat{r}_{ab}^4}{4\hat{r}_{ac}^3 \hat{r}_{bc}^3} + \frac{1}{\hat{r}_{ab}} \left(-\frac{2}{\hat{r}_{ac}} + \frac{\hat{r}_{ac}}{\hat{r}_{bc}^2} \right) \\
 & \left. \left. + \frac{1}{\hat{r}_{ab}^2} \left(3 + \frac{3\hat{r}_{ac}}{\hat{r}_{bc}} - \frac{\hat{r}_{ac}^2}{\hat{r}_{bc}^2} - \frac{\hat{r}_{ac}^3}{\hat{r}_{bc}^3} \right) + \frac{1}{\hat{r}_{ab}^3} \left(\frac{9}{2} \hat{r}_{ac} + \frac{\hat{r}_{ac}^2}{\hat{r}_{bc}} - \frac{\hat{r}_{ac}^3}{\hat{r}_{bc}^2} - \frac{\hat{r}_{ac}^4}{2\hat{r}_{bc}^3} \right) \right\} + (a \leftrightarrow b) \right]
 \end{aligned}$$

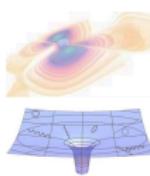


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Angular Velocity and Spin

in Newtonian mechanics and special relativity

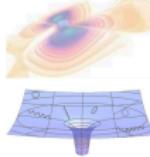
	Newton	special relativity
body-fixed frame	$x^{[i]} = \Lambda_{[i]j} x^j$	
rotational degrees of freedom ↪ supplementary condition	$\Lambda_{[k]i} \Lambda_{[k]j} = \delta_{ij}$	$\eta^{AB} \Lambda_{A\mu} \Lambda_{B\nu} = \eta_{\mu\nu}$ $\Lambda^{[i]\mu} p_\mu = 0$
Angular Velocity	$\Omega^{ij} = \Lambda_{[k]i} \frac{d\Lambda_{[k]j}}{dt}$	$\Omega^{\mu\nu} = \Lambda_A{}^\mu \frac{d\Lambda^{A\nu}}{d\tau}$
Spin (L : Lagrangian) ↪ supplementary condition	$S_{ij} = 2 \frac{\partial L}{\partial \Omega^{ij}}$	$S_{\mu\nu} = 2 \frac{\partial L}{\partial \Omega^{\mu\nu}}$ $S_{\mu\nu} p^\nu = 0$

Remark:

- Angular velocity vector is $\Omega^i = \frac{1}{2} \epsilon_{ijk} \Omega^{jk}$. Analogous for spin.



since 1558



Action approach with minimal coupling

- Gravitational field is given by a tetrad $e_{a\mu}$:

$$\Lambda_{A\mu} \Lambda^A{}_\nu = g_{\mu\nu} \quad \rightarrow \quad \Lambda_{Aa} \Lambda^A{}_b = \eta_{ab} \quad (3)$$

- Minimal coupling:

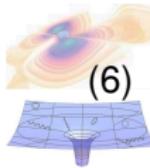
$$e^a{}_\mu e^b{}_\nu \Omega^{\mu\nu} = \Omega^{ab} = \Lambda_A{}^a \frac{D\Lambda^{Ab}}{d\tau} = \Lambda_A{}^a \left[\frac{d\Lambda^{Ab}}{d\tau} - \Lambda^A{}_c \omega_\mu{}^{cb} u^\mu \right] \quad (4)$$

- Supplementary conditions and mass-shell constraint:

$$S_{\mu\nu} p^\nu = 0, \quad \Lambda^{[i]a} e_{a\nu} p^\nu = 0, \quad p_\mu p^\mu + m^2 = 0 \quad (5)$$

- Solve constraints, supplementary and gauge conditions.
- Find variables, in which Lagrangian is of the canonical form

$$L = p_i \dot{q}^i - H$$



(6)

Canonical Structure in Detail

- In detail, the canonical Lagrangian reads: (with $\hat{\Omega}^{(i)(j)} = \hat{\Lambda}^{[k](i)}\dot{\hat{\Lambda}}^{[k](j)}$)

$$L = \frac{1}{16\pi} \int d^3x \hat{\pi}^{ij\text{TT}} \hat{h}_{ij,0}^{\text{TT}} + \hat{p}_i \dot{\hat{z}}^i + \frac{1}{2} \hat{S}_{(i)(j)} \hat{\Omega}^{(i)(j)} - H^{\text{ADM}} \quad (7)$$

- Then the equations of motion take on the form

$$\frac{dA}{dt} = \{A, H^{\text{ADM}}\} + \frac{\partial A}{\partial t} \quad (8)$$

with Poisson brackets

$$\{\hat{z}^i, \hat{p}_j\} = \delta_{ij} \quad (9)$$

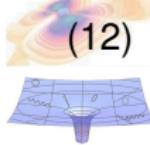
$$\{\hat{\Lambda}^{[l](j)}, \hat{S}_{(k)(l)}\} = \hat{\Lambda}^{[l](k)}\delta_{lj} - \hat{\Lambda}^{l}\delta_{kj} \quad (10)$$

$$\{\hat{S}_{(i)(j)}, \hat{S}_{(k)(l)}\} = \delta_{ik}\hat{S}_{(j)(l)} - \delta_{jk}\hat{S}_{(i)(l)} - \delta_{il}\hat{S}_{(j)(k)} + \delta_{jl}\hat{S}_{(i)(k)} \quad (11)$$



$$\{\hat{h}_{ij}^{\text{TT}}(\mathbf{x}), \hat{\pi}^{kl\text{TT}}(\mathbf{x}')\} = 16\pi\delta_{ij}^{\text{TT}kl}\delta(\mathbf{x} - \mathbf{x}') \quad (12)$$

$$\delta_{kl}^{\text{TT}ij} \equiv \text{TT-projector}$$



Canonical Variables to Linear Order in Spin

- Gauges [cf. Kibble 1963]: $e_{(0)\mu} = n_\mu$, $(e_{(i)j}) = \sqrt{(\gamma_{ij})}$, $\tau = t$
- Matter variables: compatible with SSC $\hat{S}^{\mu\nu}(p_\nu + mn_\nu) = 0$

$$z^i = \hat{z}^i - \frac{nS^i}{m - np}, \quad np = -\sqrt{m^2 + \gamma^{ij}p_i p_j}$$

$$S_{ij} = \hat{S}_{ij} - \frac{p_i n S_j}{m - np} + \frac{p_j n S_i}{m - np}, \quad n S_i = -\frac{p_k \gamma^{kj} \hat{S}_{ji}}{m}$$

$$\Lambda^{[i](j)} = \hat{\Lambda}^{[i](k)} \left(\delta_{kj} + \frac{p_{(k)} p_{(j)}}{m(m - np)} \right), \quad \gamma_{ik} \gamma_{jl} A^{kl} = \frac{1}{2} \hat{S}_{ij} + \frac{mp_{(i)} n S_{j)}}{np(m - np)}$$

$$p_i = \hat{p}_i - K_{ij} n S^j - A^{kl} e_{(j)k} e^{(j)}_{l,i} + \left(\frac{1}{2} S_{kj} + \frac{p_{(k)} n S_{j)}}{np} \right) \Gamma^{kj}_i \quad (13)$$

- Field variables:

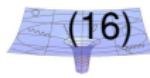
$$h_{ij}^{\text{TT}} = \hat{h}_{ij}^{\text{TT}}$$

$$\pi^{ij\text{TT}} = \hat{\pi}^{ij\text{TT}} - \delta_{kl}^{\text{TT}ij} (8\pi A^{(kl)} \delta + 16\pi B_{mn}^{kl} A^{[mn]} \delta) \quad (14)$$

$$2B_{mn}^{kl} \equiv e^{(i)}_m \frac{\partial e_{(i)n}}{\partial \gamma_{kl}} - e^{(i)}_n \frac{\partial e_{(i)m}}{\partial \gamma_{kl}}$$



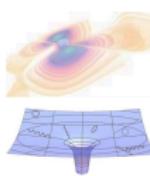
(15)



(16)

Outline

- ① Three-body interactions with spin:
NLO Spin-Orbit and NLO Spin(a)-Spin(b) Hamiltonians
 - 📄 J. Hartung and J. Steinhoff
Phys. Rev. D, submitted, arXiv:1011.1179
- ② Action approach to canonical formulation of spin in GR
 - 📄 J. Steinhoff and G. Schäfer
Europhys. Lett. **87**, 50004 (2009)
- ③ Quadrupole deformation of Neutron stars due to Spin:
NLO Spin(1)-Spin(1) Hamiltonian
 - 📄 S. Hergt, J. Steinhoff, and G. Schäfer
Class. Quant. Grav. **27**, 135007 (2010)



Quadrupole Deformation due to Spin

- Quadratic order in spin \rightarrow quadrupole deformation
- Ansatz for Dixon's quadrupole:

$$J^{\nu\rho\beta\alpha} = -3u^{[\nu}Q^{\rho][\beta}u^{\alpha]} , \quad Q_{\mu\nu} = \frac{C_Q}{m_p}S_{\mu\rho}S_{\nu}^{\rho} - \text{Trace} \quad (17)$$

- C_Q is an object dependent constant:
 - Neutron stars: $C_Q = 4 \dots 8$ [Poisson (1998); Laarakkers, Poisson (1999)]
 - Black holes: $C_Q = 1$
- Approach via effective action possible [cf. Porto, Rothstein (2008)]:

$$L_{S^2} = \underbrace{-\frac{1}{2m}R_{\mu\nu\alpha\beta}S^{\rho\mu}S^{\alpha\beta}\frac{u^\nu u_\rho}{\sqrt{-u_\sigma u^\sigma}}}_{\text{preserves supplementary conditions}} - \underbrace{\frac{1}{2}R_{\alpha\mu\beta\nu}Q^{\alpha\beta}\frac{u^\mu u^\nu}{\sqrt{-u_\sigma u^\sigma}}}_{\text{quadrupole deformation}} \quad (18)$$



$$J^{\mu\nu\alpha\beta} = -6\frac{\partial L_{S^2}}{\partial R_{\mu\nu\alpha\beta}} \quad [\text{Bailey, Israel (1975)}] \quad (19)$$



- $K_{ij,0}$ in matter action problematic for canonical formulation.

Shortcut to NLO Spin(1)-Spin(1) Hamiltonian

- \hat{p}_i -dependent part of $H_{S^2_1}^{\text{NLO}}$ follows from Poincaré Algebra.
[Hergt, Schäfer (2008)]
- Only need $\hat{p}_i = 0$ part of $H_{S^2_1}^{\text{NLO}}$.
- Only need $\hat{p}_i = 0$ part of matter energy density.
- $\hat{p}_i = 0$ part of matter energy density can be calculated from

$$\sqrt{-g} T^{\mu\nu} = \int d\tau \left[u^{(\mu} p^{\nu)} \delta_{(4)} + \left(u^{(\mu} S^{\nu)\alpha} \delta_{(4)} \right)_{||\alpha} + \frac{1}{3} R_{\alpha\beta\rho}^{(\mu} J^{\nu)\rho\beta\alpha} \delta_{(4)} - \frac{2}{3} \left(J^{\mu\alpha\beta\nu} \delta_{(4)} \right)_{||(\alpha\beta)} \right] \quad (20)$$

[Steinhoff, Pützfeld (2010)]



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$$J^{\nu\rho\beta\alpha} = -3u^{[\nu} Q^{\rho][\beta} u^{\alpha]} , \quad Q_{\mu\nu} = \frac{C_Q}{m_p} S_{\mu\rho} S_{\nu}{}^{\rho} - \text{Trace} \quad (21)$$



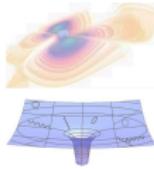
NLO Spin(1)-Spin(1) for Neutron Stars

For black holes: Steinhoff, Hergt, Schäfer (2008). See also: Porto, Rothstein (2008).

$$\begin{aligned}
 H_{S_1^2}^{\text{NLO}} = & \frac{m_2}{m_1^3 \hat{r}_{12}^3} \left[\left(\frac{15}{4} - \frac{9}{2} C_Q \right) (\hat{\mathbf{p}}_1 \cdot \hat{\mathbf{n}}_{12}) (\hat{\mathbf{S}}_1 \cdot \hat{\mathbf{n}}_{12}) (\hat{\mathbf{S}}_1 \cdot \hat{\mathbf{p}}_1) + \left(\frac{5}{4} - \frac{5}{4} C_Q \right) \hat{\mathbf{p}}_1^2 \hat{\mathbf{S}}_1^2 \right. \\
 & + \left(-\frac{9}{8} + \frac{3}{2} C_Q \right) (\hat{\mathbf{p}}_1 \cdot \hat{\mathbf{n}}_{12})^2 \hat{\mathbf{S}}_1^2 + \left(-\frac{21}{8} + \frac{9}{4} C_Q \right) \hat{\mathbf{p}}_1^2 (\hat{\mathbf{S}}_1 \cdot \hat{\mathbf{n}}_{12})^2 \\
 & \left. + \left(-\frac{5}{4} + \frac{3}{2} C_Q \right) (\hat{\mathbf{S}}_1 \cdot \hat{\mathbf{p}}_1)^2 \right] + \frac{C_Q}{m_1 m_2 \hat{r}_{12}^3} \left[\frac{9}{4} \hat{\mathbf{p}}_2^2 (\hat{\mathbf{S}}_1 \cdot \hat{\mathbf{n}}_{12})^2 - \frac{3}{4} \hat{\mathbf{p}}_2^2 \hat{\mathbf{S}}_1^2 \right] \\
 & + \frac{1}{m_1^2 \hat{r}_{12}^3} \left[\left(-\frac{3}{2} + \frac{9}{2} C_Q \right) (\hat{\mathbf{p}}_2 \cdot \hat{\mathbf{n}}_{12}) (\hat{\mathbf{S}}_1 \cdot \hat{\mathbf{n}}_{12}) (\hat{\mathbf{S}}_1 \cdot \hat{\mathbf{p}}_1) \right. \\
 & + \left(-3 + \frac{3}{2} C_Q \right) (\hat{\mathbf{p}}_1 \cdot \hat{\mathbf{n}}_{12}) (\hat{\mathbf{S}}_1 \cdot \hat{\mathbf{n}}_{12}) (\hat{\mathbf{S}}_1 \cdot \hat{\mathbf{p}}_2) + \left(-\frac{3}{2} + \frac{9}{4} C_Q \right) (\hat{\mathbf{p}}_1 \cdot \hat{\mathbf{p}}_2) \hat{\mathbf{S}}_1^2 \\
 & + \left(\frac{3}{2} - \frac{3}{4} C_Q \right) (\hat{\mathbf{p}}_1 \cdot \hat{\mathbf{n}}_{12}) (\hat{\mathbf{p}}_2 \cdot \hat{\mathbf{n}}_{12}) \hat{\mathbf{S}}_1^2 + \left(3 - \frac{21}{4} C_Q \right) (\hat{\mathbf{p}}_1 \cdot \hat{\mathbf{p}}_2) (\hat{\mathbf{S}}_1 \cdot \hat{\mathbf{n}}_{12})^2 \\
 & \left. - \frac{15}{4} C_Q (\hat{\mathbf{p}}_1 \cdot \hat{\mathbf{n}}_{12}) (\hat{\mathbf{p}}_2 \cdot \hat{\mathbf{n}}_{12}) (\hat{\mathbf{S}}_1 \cdot \hat{\mathbf{n}}_{12})^2 + \left(\frac{3}{2} - \frac{3}{2} C_Q \right) (\hat{\mathbf{S}}_1 \cdot \hat{\mathbf{p}}_1) (\hat{\mathbf{S}}_1 \cdot \hat{\mathbf{p}}_2) \right] \\
 & + \frac{m_2}{\hat{r}_{12}^4} \left[\left(2 + \frac{1}{2} C_Q \right) \hat{\mathbf{S}}_1^2 - \left(3 + \frac{3}{2} C_Q \right) (\hat{\mathbf{S}}_1 \cdot \hat{\mathbf{n}}_{12})^2 \right] \\
 & + \frac{m_2^2}{m_1 \hat{r}_{12}^4} \left[(1 + 2C_Q) \hat{\mathbf{S}}_1^2 - (1 + 6C_Q) (\hat{\mathbf{S}}_1 \cdot \hat{\mathbf{n}}_{12})^2 \right]
 \end{aligned}$$



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Thank you for your attention

and the German Research Foundation **DFG** for support

