

Canonical formulation of spinning objects in General Relativity

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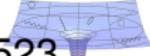
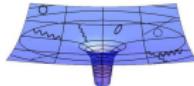
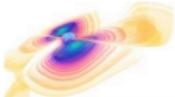
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DFG: SFB/TR7 “Gravitational Wave Astronomy” and GRK 1523

Outline

1 Introduction

2 Action Approach

3 Order-by-Order Construction

4 Results linear in spin

5 Higher orders in spin



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(3+1)-Decomposition

- Normal vector:

$$n_\mu = (-N, 0, 0, 0), \quad n^\mu = \frac{1}{N}(1, -N^i), \quad n_\mu n^\mu = -1$$

- Projector:

$$\gamma^{\mu\nu} = g^{\mu\nu} + n^\mu n^\nu = \begin{pmatrix} 0 & 0 \\ 0 & \gamma^{ij} \end{pmatrix}, \quad g_{ij} = \gamma_{ij}, \quad \gamma_{ik}\gamma^{kj} = \delta_{ij}$$

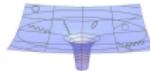
- Extrinsic curvature:

$$K_{ij} \equiv -n_{(i||j)}$$

$$\pi^{ij} = -\sqrt{\gamma}(\gamma^{ik}\gamma^{lj} - \gamma^{ij}\gamma^{kl})K_{kl}$$



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Point-Mass Action in ADM Form

$$W = \int dt \mathbf{p}_i \dot{\mathbf{z}}^i + \int d^4x \left[\frac{1}{16\pi} \pi^{ij} \gamma_{ij,0} - N \mathcal{H} + N^i \mathcal{H}_i + (\text{st}) \right]$$

- Constraint equations:

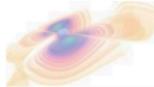
$$0 = \mathcal{H} \equiv -\frac{1}{16\pi\sqrt{\gamma}} \left[\gamma R + \frac{1}{2} \left(\gamma_{ij} \pi^{ij} \right)^2 - \gamma_{ij} \gamma_{kl} \pi^{ik} \pi^{jl} \right] + \mathcal{H}^{\text{matter}}$$

$$0 = \mathcal{H}_i \equiv \frac{1}{8\pi} \gamma_{ij} \pi^{jk}_{;k} + \mathcal{H}_i^{\text{matter}}$$

- Source terms are related to the stress-energy tensor $T^{\mu\nu}$:

$$\mathcal{H}^{\text{matter}} \equiv \sqrt{\gamma} T_{\mu\nu} n^\mu n^\nu = \sqrt{m^2 + \gamma^{ij} p_i p_j} \delta$$

$$\mathcal{H}_i^{\text{matter}} \equiv -\sqrt{\gamma} T_{i\nu} n^\nu = p_i \delta$$



ADM Canonical Formalism

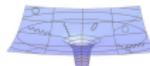
- Hamiltonian without gauge fixing:

$$H[x_a^i, p_{ai}, \gamma_{ij}, \pi^{ij}] = \int d^3\mathbf{x} (N\mathcal{H} - N^i \mathcal{H}_i) + E[\gamma_{ij}]$$
$$E[\gamma_{ij}] = \frac{1}{16\pi} \oint d^2 s_i (\gamma_{ij,j} - \gamma_{jj,i})$$

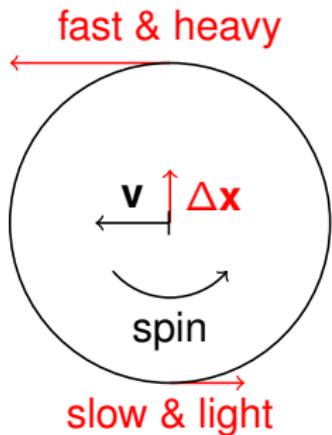
- ADM Hamiltonian (Hamiltonian in ADMTT gauge)
 $\hat{=}$ ADM Energy depending on canonical variables:

$$H_{\text{ADM}} = E[x_a^i, p_{ai}, h_{ij}^{\text{TT}}, \pi_{\text{TT}}^{ij}] = -\frac{1}{16\pi} \int d^3\mathbf{x} \Delta \phi$$
$$\gamma_{ij} = \left(1 + \frac{1}{8}\phi\right)^4 \delta_{ij} + h_{ij}^{\text{TT}}, \quad \pi^{ii} = 0$$

- Matter only Hamiltonian: Elimination of h_{ij}^{TT} and π_{TT}^{ij} .



Spin in Special Relativity



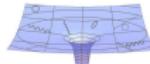
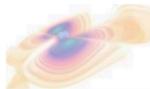
- Different mass centers.
- Spin is a 4-tensor $S^{\mu\nu}$:
 - Spin is $S^{ij} = \epsilon^{ijk} S_k$.
 - Mass dipole related to S^{i0} .
- Need spin supplementary condition:
 - Møller SSC: $\tilde{S}^{\mu 0} = 0$
 - Covariant SSC: $S^{\mu\nu} p_\nu = 0$
 - **Newton-Wigner (canonical) SSC:**
 $m\hat{S}^{\mu 0} + \hat{S}^{\mu\nu} p_\nu = 0$
- In covariant SSC, with position **z**:
- In Newton-Wigner SSC:

$$\{z^i(t), z^j(t)\} = \frac{S^{ij}}{m^2} - \frac{p^i S^{0j} - p^j S^{0i}}{m^2 p^0}, \quad \dots$$

$$\{\hat{z}^i(t), p_j(t)\} = \delta_{ij}, \quad \{\hat{S}_i(t), \hat{S}_j(t)\} = \epsilon_{ijk} \hat{S}_k(t)$$



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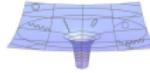
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Non-Relativistic Spherical Top

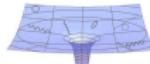
- Body fixed coordinates \tilde{x}_a^i : $x_a^i(t) = R_{ji}(t)\tilde{x}_a^j$, $R_{ki}R_{kj} = \delta_{ij}$
- Independent angle variables: $R_{ij} = R_{ij}(\phi, \psi, \theta)$
- Angular velocity tensor: $\Omega^{ij} = \varepsilon_{ijk}\omega^k = R_{ki}\dot{R}_{kj}$
- Lagrangian: $L = \frac{1}{4}J\Omega^{ij}\Omega^{ij}$
- Spin tensor: $S_{ij} = 2\frac{\partial L}{\partial \Omega^{ij}} = J\Omega^{ij}$
- Legendre transformed:

$$L = \frac{1}{2}S_{ij}\Omega^{ij} - H[R_{ij}, S_{ij}], \quad H = \frac{1}{4J}S_{ij}S_{ij}$$

- Trick: Use $\delta\theta_{ij} = -\delta\theta_{ji} = R_{ki}\delta R_{kj}$ as independent variations.
- Usual Poisson brackets.



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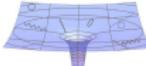


Relativistic Spherical Top

- No rigid bodies.
- Mathematical abstraction: Top is
 - Worldline with Lorentz-matrix $\Lambda_{A\mu}$. $\eta^{AB}\Lambda_{A\mu}\Lambda_{B\nu} = \eta_{\mu\nu}$
 - $\Lambda_{A\mu}$ is pure rotation in rest-frame: $\Lambda_{A\mu} = \begin{pmatrix} -1 & 0 \\ 0 & R_{ij} \end{pmatrix}$
- Equivalent description: $\Lambda_{[0]\mu} = p_\mu/m$ or $\Lambda^{[i]\mu} p_\mu = 0$
- Angular velocity tensor: $\Omega^{\mu\nu} = \Lambda_A{}^\mu \frac{d\Lambda^{Av}}{d\tau}$
- Spin tensor: $S_{\mu\nu} = 2 \frac{\partial L}{\partial \Omega^{\mu\nu}}$
- Associated SSC: $S_{\mu\nu} p^\mu = 0$



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Minimal Coupling

- Problem with metric variation: (also $\Lambda_{A\mu} = e_{A\mu}$)

$$\Lambda_{A\mu} \Lambda^A{}_v = g_{\mu v} \leftrightarrow [\gamma_\mu, \gamma_v]_A = 2g_{\mu v}$$

- Variate Λ^{Aa} and tetrad $e_{a\mu}$, $\Lambda^A{}_\mu = \Lambda^{Aa} e_{a\mu}$:

$$\Lambda_{Aa} \Lambda^A{}_b = \eta_{ab} \leftrightarrow [\gamma_a, \gamma_b]_A = 2\eta_{ab}$$

- Matter Lagrangian density and constraints:

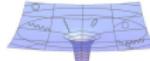
$$\mathcal{L}_M = \int d\tau \left[p_\mu u^\mu + \frac{1}{2} S_{ab} \Omega^{ab} \right] \delta_{(4)}$$

$$\Omega^{ab} = \Lambda_A{}^a \frac{d\Lambda^{Ab}}{d\tau} - \omega_\mu{}^{ab} u^\mu$$

$$S_{ab} p^b = 0, \quad \Lambda^{[i]a} p_a = 0, \quad p_\mu p^\mu + m^2 = 0$$



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Result of Variation

- Approximated linear in spin.
- Field equations with stress-energy tensor (Mathisson, Tulczyjew):

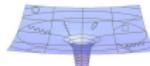
$$\sqrt{-g} T^{\mu\nu} = \int d\tau \left[mu^\mu u^\nu \delta_{(4)} - (S^{\alpha(\mu} u^{\nu)} \delta_{(4)})_{||\alpha} \right]$$

- EOM (Mathisson, Papapetrou):

$$\frac{DS^{\mu\nu}}{d\tau} = 0, \quad \frac{Dp_\mu}{d\tau} = \frac{1}{2} S^{\lambda\nu} u^\gamma R_{\mu\gamma\nu\lambda}$$

- Derivation: Evaluate $T^{\mu\nu}_{||\nu} = 0$ with ansatz

$$\sqrt{-g} T^{\mu\nu} = \int d\tau \left[t^{\mu\nu} \delta_{(4)} + (t^{\mu\nu\alpha} \delta_{(4)})_{||\alpha} \right]$$



Reduction of the Matter Part

- Solve matter constraints, Schwinger time gauge $e_{(0)\mu} = n_\mu$, $\tau = t$.
- Variable redefinitions:

$$z^i = \hat{z}^i - \frac{nS^i}{m-np}, \quad np = -\sqrt{m^2 + \gamma^{ij} p_i p_j}$$

$$S_{ij} = \hat{S}_{ij} - \frac{p_i n S_j}{m-np} + \frac{p_j n S_i}{m-np}, \quad n S_i = -\frac{p_k \gamma^{kj} \hat{S}_{ji}}{m}$$

$$\Lambda^{[i](j)} = \hat{\Lambda}^{[i](k)} \left(\delta_{kj} + \frac{p_{(k)} p^{(j)}}{m(m-np)} \right), \quad \gamma_{ik} \gamma_{jl} A^{kl} = \frac{1}{2} \hat{S}_{ij} + \frac{mp_{(i} n S_{j)}}{np(m-np)}$$

$$\hat{p}_i = p_i + K_{ij} n S^j + A^{kl} e_{(j)k} e_{l,i} - \left(\frac{1}{2} S_{kj} + \frac{p_{(k} n S_{j)}}{np} \right) \Gamma^{kj}_i$$

- Matter Lagrangian density now, with $\hat{\Omega}^{(i)(j)} = \hat{\Lambda}_{[k]}^{(i)} \dot{\hat{\Lambda}}^{[k](j)}$,

$$\mathcal{L}_M = A^{ij} e_{(k)i} e_{j,0}^{(k)} \hat{\delta} + \hat{p}_i \dot{\hat{z}}^i \hat{\delta} + \frac{1}{2} \hat{S}_{(i)(j)} \hat{\Omega}^{(i)(j)} \hat{\delta} - N \mathcal{H}^{\text{matter}} + N^i \mathcal{H}_i^{\text{matter}}$$



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ADM Formalism with Spin

- Legendre transformation for gravitational field.
- Spatial symmetric gauge (Kibble 1963): $e_{(i)j} = e_{ij} = e_{ji}$

$$e_{ij} e_{jk} = \gamma_{ik} \quad \Rightarrow \quad (e_{ij}) = \sqrt{(\gamma_{ij})}$$

- Action:

$$W = \frac{1}{16\pi} \int d^4x \hat{\pi}^{ij} \gamma_{ij,0} + \int dt \left[\hat{p}_i \dot{\hat{z}}^i + \frac{1}{2} \hat{S}_{(i)(j)} \hat{\Omega}^{(i)(j)} - H \right]$$

$$H = \int d^3x (N \mathcal{H} - N^i \mathcal{H}_i) + E[\gamma_{ij}]$$

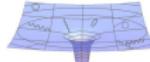
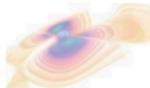
- Canonical field momentum:

$$\hat{\pi}^{ij} = \pi^{ij} + 8\pi A^{(ij)} \hat{\delta} + 16\pi B_{kl}^{ij} A^{[kl]} \hat{\delta}$$

$$e_{k[i} e_{j]k,0} = B_{ij}^{kl} \gamma_{kl,0}$$



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Full Reduction

- Solve field constraints in ADMTT gauge.
- Fully reduced action:

$$W = \frac{1}{16\pi} \int d^4x \hat{\pi}_{\text{TT}}^{ij} h_{ij,0}^{\text{TT}} + \int dt \left[\hat{p}_i \dot{\hat{z}}^i + \frac{1}{2} \hat{S}_{(i)(j)} \hat{\Omega}^{(i)(j)} - H_{\text{ADM}} \right]$$
$$H_{\text{ADM}} = E[\hat{z}^i, \hat{p}_i, \hat{S}_{(i)(j)}, h_{ij}^{\text{TT}}, \hat{\pi}^{ij\text{TT}}]$$

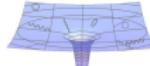
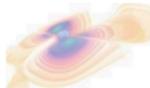
- Fundamental equal-time Poisson brackets:

$$\{\hat{z}_a^i, \hat{p}_{aj}\} = \delta_{ij}, \quad \{\hat{S}_{a(i)}, \hat{S}_{a(j)}\} = \epsilon_{ijk} \hat{S}_{a(k)}$$
$$\{h_{ij}^{\text{TT}}(\mathbf{x}, t), \hat{\pi}_{\text{TT}}^{kl}(\mathbf{x}', t)\} = 16\pi \delta_{ij}^{\text{TT}kl} \delta(\mathbf{x} - \mathbf{x}')$$

- Valid to all orders linear in spin.



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Conserved Quantities

- Energy: $E = H_{\text{ADM}}$
- Total linear and angular momentum:

$$P_i = \sum_a \hat{p}_{ai} - \frac{1}{16\pi} \int d^3x \hat{\pi}_{\text{TT}}^{kl} h_{kl,i}^{\text{TT}}$$
$$J_{ij} = \sum_a (\hat{z}_a^i \hat{p}_{aj} - \hat{z}_a^j \hat{p}_{ai}) + \sum_a \hat{S}_{a(i)(j)}$$
$$- \frac{1}{16\pi} \int d^3x (x^i \hat{\pi}_{\text{TT}}^{kl} h_{kl,j}^{\text{TT}} - x^j \hat{\pi}_{\text{TT}}^{kl} h_{kl,i}^{\text{TT}})$$
$$- \frac{1}{16\pi} \int d^3x 2(\hat{\pi}_{\text{TT}}^{ik} h_{kj}^{\text{TT}} - \hat{\pi}_{\text{TT}}^{jk} h_{ki}^{\text{TT}})$$

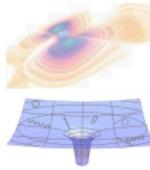
- Boost: $J^{i0} \equiv K^i \equiv G^i - t P^i$

With center-of-mass vector:

$$G^i = -\frac{1}{16\pi} \int d^3x x^i \Delta\phi$$



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Stress-Energy-Tensor Algebra (Minkowski)

$$\{\mathcal{H}^m(x), \mathcal{H}^m(x')\} = -\mathcal{H}_i^m(x)\delta_{xx',i} - \mathcal{H}_i^m(x')\delta_{xx',i}$$

$$\{\mathcal{H}_i^m(x), \mathcal{H}^m(x')\} = -\mathcal{H}^m(x)\delta_{xx',i} - T_{ij}(x')\delta_{xx',j}$$

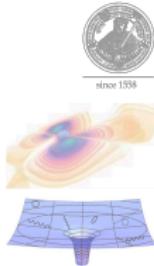
$$\{\mathcal{H}_i^m(x), \mathcal{H}_j^m(x')\} = -\mathcal{H}_j^m(x)\delta_{xx',i} - \mathcal{H}_i^m(x')\delta_{xx',j} + \partial_n \partial'_q [h_{injq}(x)\delta_{xx'}]$$

$$h_{injq}(x) = \left[-\hat{S}_{q)(n}\mathcal{P}_i)(j} - \delta^{kl} \frac{p_k \hat{S}_{l(n}\mathcal{P}_i)(j} p_q)}{(np)(m-np)} + \delta^{kl} \frac{p_k \hat{S}_{l(q}\mathcal{P}_j)(i} p_n)}{(np)(m-np)} \right] \delta$$

$$\mathcal{P}_{ij} \equiv \delta_{ij} - \frac{p_i p_j}{(np)^2}$$

$$\mathcal{H}^m(x) = T^{00}, \quad \mathcal{H}_i^m(x) = T^{0i}$$

- Local version of the Poincaré algebra.
- Minkowski limit of the gravitational constraint algebra.
- Dirac field also has $h_{injq}(x) \neq 0$.

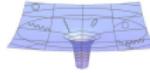


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Order-by-Order Construction

- Hamiltonian $\hat{=}$ energy depending on canonical variables.
- Need to find canonical variables!
- Symmetries of the action:

$$P_i = \sum_a \hat{p}_{ai} + P_i^{\text{field}}, \quad J_i = \sum_a \left[\varepsilon_{ijk} \hat{z}_a^j \hat{p}_{ak} + S_{a(i)} \right] + J_i^{\text{field}}$$

- Global Poincaré algebra:

$$\{P_i, P_j\} = 0, \quad \{P_i, H\} = 0, \quad \{J_i, H\} = 0$$

$$\{J_i, P_j\} = \varepsilon_{ijk} P_k, \quad \{J_i, J_j\} = \varepsilon_{ijk} J_k, \quad \{J_i, G_j\} = \varepsilon_{ijk} G_k$$

$$\{G_i, P_j\} = H \delta_{ij}, \quad \{G_i, H\} = P_i, \quad \{G_i, G_j\} = -\varepsilon_{ijk} J_k$$



Canonical Variables at 2PN

- Calculate $\mathcal{H}_i^{\text{matter}}$:

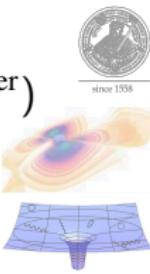
$$\mathcal{H}_i^{\text{matter}} = -\sqrt{\gamma} T_{i\nu} n^\nu$$

- Define canonical momentum \hat{p}_i as:

$$\hat{p}_i = \int d^3\mathbf{x} \mathcal{H}_i^{\text{matter}}$$

- Define spin $\hat{S}_{ij} = e_{i(k)} e_{j(l)} \epsilon_{klm} \hat{S}_{(m)}$ such that $\hat{\mathbf{S}}^2 = \text{const.}$
- $\hat{\mathbf{z}}$ and $e_{i(k)}$ fixed by

$$J_{ij} = \hat{z}^i \hat{p}_j - \hat{z}^j \hat{p}_i + \epsilon_{ijm} \hat{S}_{(m)} = \int d^3\mathbf{x} (x^i \mathcal{H}_j^{\text{matter}} - x^j \mathcal{H}_i^{\text{matter}})$$



Outline

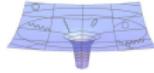
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Leading-Order (LO)

- LO Spin-Orbit Hamiltonian:

$$H_{\text{SO}}^{\text{LO}} = \sum_a \sum_{b \neq a} \frac{1}{\hat{r}_{ab}^2} (\hat{\mathbf{S}}_a \times \hat{\mathbf{n}}_{ab}) \cdot \left[\frac{3m_b}{2m_a} \hat{\mathbf{p}}_a - 2\hat{\mathbf{p}}_b \right]$$

- LO Spin₁-Spin₂ Hamiltonian:

$$H_{S_1 S_2}^{\text{LO}} = \sum_a \sum_{b \neq a} \frac{1}{2\hat{r}_{ab}^3} \left[3(\hat{\mathbf{S}}_a \cdot \hat{\mathbf{n}}_{ab})(\hat{\mathbf{S}}_b \cdot \hat{\mathbf{n}}_{ab}) - (\hat{\mathbf{S}}_a \cdot \hat{\mathbf{S}}_b) \right]$$

- Center of mass vector:

$$\mathbf{G}_{\text{SO}}^{\text{LO}} = \sum_a \frac{1}{2m_a} (\hat{\mathbf{p}}_a \times \hat{\mathbf{S}}_a), \quad \mathbf{G}_{S_1 S_2}^{\text{LO}} = 0$$



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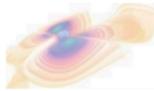
NLO Spin-Orbit

First derived: Damour, Jaranowski, and Schäfer (2008)

$$\begin{aligned} H_{\text{SO}}^{\text{NLO}} = & - \frac{((\hat{\mathbf{p}}_1 \times \hat{\mathbf{S}}_1) \cdot \hat{\mathbf{n}}_{12})}{\hat{r}_{12}^2} \left[\frac{5m_2 \hat{\mathbf{p}}_1^2}{8m_1^3} + \frac{3(\hat{\mathbf{p}}_1 \cdot \hat{\mathbf{p}}_2)}{4m_1^2} - \frac{3\hat{\mathbf{p}}_2^2}{4m_1 m_2} \right. \\ & \quad \left. + \frac{3(\hat{\mathbf{p}}_1 \cdot \hat{\mathbf{n}}_{12})(\hat{\mathbf{p}}_2 \cdot \hat{\mathbf{n}}_{12})}{4m_1^2} + \frac{3(\hat{\mathbf{p}}_2 \cdot \hat{\mathbf{n}}_{12})^2}{2m_1 m_2} \right] \\ & + \frac{((\hat{\mathbf{p}}_2 \times \hat{\mathbf{S}}_1) \cdot \hat{\mathbf{n}}_{12})}{\hat{r}_{12}^2} \left[\frac{(\hat{\mathbf{p}}_1 \cdot \hat{\mathbf{p}}_2)}{m_1 m_2} + \frac{3(\hat{\mathbf{p}}_1 \cdot \hat{\mathbf{n}}_{12})(\hat{\mathbf{p}}_2 \cdot \hat{\mathbf{n}}_{12})}{m_1 m_2} \right] \\ & + \frac{((\hat{\mathbf{p}}_1 \times \hat{\mathbf{S}}_1) \cdot \hat{\mathbf{p}}_2)}{\hat{r}_{12}^2} \left[\frac{2(\hat{\mathbf{p}}_2 \cdot \hat{\mathbf{n}}_{12})}{m_1 m_2} - \frac{3(\hat{\mathbf{p}}_1 \cdot \hat{\mathbf{n}}_{12})}{4m_1^2} \right] \\ & - \frac{((\hat{\mathbf{p}}_1 \times \hat{\mathbf{S}}_1) \cdot \hat{\mathbf{n}}_{12})}{\hat{r}_{12}^3} \left[\frac{11m_2}{2} + \frac{5m_2^2}{m_1} \right] \\ & + \frac{((\hat{\mathbf{p}}_2 \times \hat{\mathbf{S}}_1) \cdot \hat{\mathbf{n}}_{12})}{\hat{r}_{12}^3} \left[6m_1 + \frac{15m_2}{2} \right] + (1 \leftrightarrow 2) \end{aligned}$$



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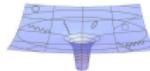
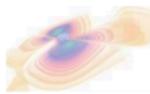
NLO Spin₁-Spin₂

Partial result: Porto and Rothstein (2006)

$$\begin{aligned} H_{S_1 S_2}^{\text{NLO}} = & \frac{1}{2m_1 m_2 \hat{r}_{12}^3} [\frac{3}{2} ((\hat{\mathbf{p}}_1 \times \hat{\mathbf{S}}_1) \cdot \hat{\mathbf{n}}_{12}) ((\hat{\mathbf{p}}_2 \times \hat{\mathbf{S}}_2) \cdot \hat{\mathbf{n}}_{12}) + \frac{1}{2} (\hat{\mathbf{S}}_1 \cdot \hat{\mathbf{S}}_2) (\hat{\mathbf{p}}_1 \cdot \hat{\mathbf{p}}_2) \\ & + 6((\hat{\mathbf{p}}_2 \times \hat{\mathbf{S}}_1) \cdot \hat{\mathbf{n}}_{12}) ((\hat{\mathbf{p}}_1 \times \hat{\mathbf{S}}_2) \cdot \hat{\mathbf{n}}_{12}) - \frac{1}{2} (\hat{\mathbf{S}}_1 \cdot \hat{\mathbf{p}}_2) (\hat{\mathbf{S}}_2 \cdot \hat{\mathbf{p}}_1) \\ & - 15(\hat{\mathbf{S}}_1 \cdot \hat{\mathbf{n}}_{12}) (\hat{\mathbf{S}}_2 \cdot \hat{\mathbf{n}}_{12}) (\hat{\mathbf{p}}_1 \cdot \hat{\mathbf{n}}_{12}) (\hat{\mathbf{p}}_2 \cdot \hat{\mathbf{n}}_{12}) + (\hat{\mathbf{S}}_1 \cdot \hat{\mathbf{p}}_1) (\hat{\mathbf{S}}_2 \cdot \hat{\mathbf{p}}_2) \\ & - 3(\hat{\mathbf{S}}_1 \cdot \hat{\mathbf{n}}_{12}) (\hat{\mathbf{S}}_2 \cdot \hat{\mathbf{n}}_{12}) (\hat{\mathbf{p}}_1 \cdot \hat{\mathbf{p}}_2) + 3(\hat{\mathbf{S}}_1 \cdot \hat{\mathbf{p}}_2) (\hat{\mathbf{S}}_2 \cdot \hat{\mathbf{n}}_{12}) (\hat{\mathbf{p}}_1 \cdot \hat{\mathbf{n}}_{12}) \\ & + 3(\hat{\mathbf{S}}_2 \cdot \hat{\mathbf{p}}_1) (\hat{\mathbf{S}}_1 \cdot \hat{\mathbf{n}}_{12}) (\hat{\mathbf{p}}_2 \cdot \hat{\mathbf{n}}_{12}) + 3(\hat{\mathbf{S}}_1 \cdot \hat{\mathbf{p}}_1) (\hat{\mathbf{S}}_2 \cdot \hat{\mathbf{n}}_{12}) (\hat{\mathbf{p}}_2 \cdot \hat{\mathbf{n}}_{12}) \\ & + 3(\hat{\mathbf{S}}_2 \cdot \hat{\mathbf{p}}_2) (\hat{\mathbf{S}}_1 \cdot \hat{\mathbf{n}}_{12}) (\hat{\mathbf{p}}_1 \cdot \hat{\mathbf{n}}_{12}) - 3(\hat{\mathbf{S}}_1 \cdot \hat{\mathbf{S}}_2) (\hat{\mathbf{p}}_1 \cdot \hat{\mathbf{n}}_{12}) (\hat{\mathbf{p}}_2 \cdot \hat{\mathbf{n}}_{12})] \\ & + \frac{3}{2m_1^2 \hat{r}_{12}^3} [- ((\hat{\mathbf{p}}_1 \times \hat{\mathbf{S}}_1) \cdot \hat{\mathbf{n}}_{12}) ((\hat{\mathbf{p}}_1 \times \hat{\mathbf{S}}_2) \cdot \hat{\mathbf{n}}_{12}) \\ & \quad + (\hat{\mathbf{S}}_1 \cdot \hat{\mathbf{S}}_2) (\hat{\mathbf{p}}_1 \cdot \hat{\mathbf{n}}_{12})^2 - (\hat{\mathbf{S}}_1 \cdot \hat{\mathbf{n}}_{12}) (\hat{\mathbf{S}}_2 \cdot \hat{\mathbf{p}}_1) (\hat{\mathbf{p}}_1 \cdot \hat{\mathbf{n}}_{12})] \\ & + \frac{3}{2m_2^2 \hat{r}_{12}^3} [- ((\hat{\mathbf{p}}_2 \times \hat{\mathbf{S}}_2) \cdot \hat{\mathbf{n}}_{12}) ((\hat{\mathbf{p}}_2 \times \hat{\mathbf{S}}_1) \cdot \hat{\mathbf{n}}_{12}) \\ & \quad + (\hat{\mathbf{S}}_1 \cdot \hat{\mathbf{S}}_2) (\hat{\mathbf{p}}_2 \cdot \hat{\mathbf{n}}_{12})^2 - (\hat{\mathbf{S}}_2 \cdot \hat{\mathbf{n}}_{12}) (\hat{\mathbf{S}}_1 \cdot \hat{\mathbf{p}}_2) (\hat{\mathbf{p}}_2 \cdot \hat{\mathbf{n}}_{12})] \\ & + \frac{6(m_1 + m_2)}{\hat{r}_{12}^4} [(\hat{\mathbf{S}}_1 \cdot \hat{\mathbf{S}}_2) - 2(\hat{\mathbf{S}}_1 \cdot \hat{\mathbf{n}}_{12}) (\hat{\mathbf{S}}_2 \cdot \hat{\mathbf{n}}_{12})] \end{aligned}$$



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NLO Center of Mass

$$\begin{aligned}\mathbf{G}_{\text{SO}}^{\text{NLO}} = & - \sum_a \frac{\hat{\mathbf{p}}_a^2}{8m_a^3} (\hat{\mathbf{p}}_a \times \hat{\mathbf{S}}_a) \\ & + \sum_a \sum_{b \neq a} \frac{m_b}{4m_a \hat{r}_{ab}} \left[((\hat{\mathbf{p}}_a \times \hat{\mathbf{S}}_a) \cdot \hat{\mathbf{n}}_{ab}) \frac{5\hat{\mathbf{z}}_a + \hat{\mathbf{z}}_b}{\hat{r}_{ab}} - 5(\hat{\mathbf{p}}_a \times \hat{\mathbf{S}}_a) \right] \\ & + \sum_a \sum_{b \neq a} \frac{1}{\hat{r}_{ab}} \left[\frac{3}{2} (\hat{\mathbf{p}}_b \times \hat{\mathbf{S}}_a) - \frac{1}{2} (\hat{\mathbf{n}}_{ab} \times \hat{\mathbf{S}}_a) (\hat{\mathbf{p}}_b \cdot \hat{\mathbf{n}}_{ab}) \right. \\ & \quad \left. - ((\hat{\mathbf{p}}_b \times \hat{\mathbf{S}}_a) \cdot \hat{\mathbf{n}}_{ab}) \frac{\hat{\mathbf{z}}_a + \hat{\mathbf{z}}_b}{\hat{r}_{ab}} \right]\end{aligned}$$

$$\mathbf{G}_{\text{SS}}^{\text{NLO}} = \frac{1}{2} \sum_a \sum_{b \neq a} \left\{ \left[3(\hat{\mathbf{S}}_a \cdot \hat{\mathbf{n}}_{ab})(\hat{\mathbf{S}}_b \cdot \hat{\mathbf{n}}_{ab}) - (\hat{\mathbf{S}}_a \cdot \hat{\mathbf{S}}_b) \right] \frac{\hat{\mathbf{z}}_a}{\hat{r}_{ab}^3} + (\hat{\mathbf{S}}_b \cdot \hat{\mathbf{n}}_{ab}) \frac{\hat{\mathbf{S}}_a}{\hat{r}_{ab}^2} \right\}$$

⇒ Poincaré algebra is fulfilled.

Outline

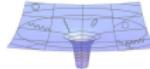
1 Introduction

2 Action Approach

3 Order-by-Order Construction

4 Results linear in spin

5 Higher orders in spin



The Stress-Energy Tensor with Quadrupole

- Ansatz:

$$\sqrt{-g} T^{\mu\nu} = \int d\tau \left[t^{\mu\nu} \delta_{(4)} + (t^{\mu\nu\alpha} \delta_{(4)})_{||\alpha} + (\textcolor{red}{t^{\mu\nu\alpha\beta}} \delta_{(4)})_{||\alpha\beta} \right]$$

- Evaluate $T_{||\nu}^{\mu\nu} = 0$:

$$\frac{D(S^{\mu\nu})}{D\tau} = 2p^{[\mu} u^{\nu]} + \frac{4}{3} R_{\rho\beta\alpha}^{[\mu} J^{\nu]\rho\beta\alpha}$$

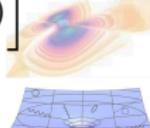
$$\frac{Dp_\mu}{D\tau} = -\frac{1}{2} R_{\mu\rho\beta\alpha} u^\rho S^{\beta\alpha} - \frac{1}{6} R_{\nu\rho\beta\alpha;\mu} J^{\nu\rho\beta\alpha}$$

$$\begin{aligned} \sqrt{-g} T^{\mu\nu} = \int d\tau & \left[u^{(\mu} p^{\nu)} \delta_{(4)} + \frac{1}{3} R_{\rho\beta\alpha}^{(\mu} J^{\nu)\rho\beta\alpha} \delta_{(4)} \right. \\ & \left. + (u^{(\mu} S^{\nu)\alpha} \delta_{(4)})_{||\alpha} - \frac{2}{3} (\textcolor{red}{J^{\mu\alpha\beta\nu}} \delta_{(4)})_{||(\alpha\beta)} \right] \end{aligned}$$

- Mass quadrupole $I^{\mu\nu}$: $\textcolor{red}{J^{\nu\rho\beta\alpha}} = -3u^{[\nu} I^{\rho][\beta} u^{\alpha]}$
- Spin-squared ansatz for $I^{\mu\nu}$.



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Hamiltonians from Kerr Metric and Poincaré Algebra

Hergt and Schäfer

Full Hamiltonian relevant for Poincaré algebra up to 2PN:

$$H = H_N + H_{1PN} + H_{2PN} + H_{SO}^{1PN} + H_{SO}^{2PN} + H_{S^2} + H_{S^3 p} + H_{S^2 p^2} + H_{S^4}$$

- Source terms in canonical variables sufficient for $H_{S_2^2 S_1 p_1}$, $H_{S_2^3 p_1}$, $H_{S_1^3 p_2}$, $H_{S_1^2 S_2 p_2}$, $H_{S_1^2 S_2^2}$, $H_{S_1 S_2^3}$, and $H_{S_2 S_1^3}$ were obtained from the Kerr-metric in ADM coordinates.
- Ansatzes for $H_{S_1^2 p^2}$, $H_{S_2^2 p^2}$, $H_{S_1^3 p_1}$, $H_{S_2^3 p_2}$, $H_{S_1^2 S_2 p_1}$, $H_{S_2^2 S_1 p_2}$, $H_{S_1^4}$, and $H_{S_2^4}$ are **fixed up to canonical transformation** by $\{G_i, H\} = P_i$.
- The \hat{p}_i -independent part of the Hamilton constraint is needed to fix these remaining degrees of freedom and to get $H_{G^2 S^2}$.



Ansatz for $\hat{p}_i = 0$ Source Terms

$$\begin{aligned}\mathcal{H}_{\mathbf{S}_1^2, \hat{p}_i=0}^{\text{matter}} &= \frac{c_1}{m_1} \left(Q_1^{ij} \hat{\delta}_1 \right)_{;ij} + \frac{c_2}{m_1} R_{ij} Q_1^{ij} \hat{\delta}_1 + \frac{c_3}{m_1} \hat{\mathbf{S}}_1^2 \left(\gamma^{ij} \delta_1 \right)_{;ij} + \frac{c_4}{m_1} R \hat{\mathbf{S}}_1^2 \hat{\delta}_1 \\ &\quad + \frac{1}{8m_1} g_{mn} \gamma^{pj} \gamma^{ql} \gamma^{mi}_{,p} \gamma^{nk}_{,q} \hat{S}_{1ij} \hat{S}_{1kl} \hat{\delta}_1 + \frac{1}{4m_1} \left(\gamma^{ij} \gamma^{mn} \gamma^{kl}_{,m} \hat{S}_{1ln} \hat{S}_{1jk} \hat{\delta}_1 \right)_{,i}\end{aligned}$$

$$Q_1^{ij} = \gamma^{ik} \gamma^{jl} \gamma^{mn} \hat{S}_{1km} \hat{S}_{1nl} - \frac{2}{3} \gamma^{ij} \hat{\mathbf{S}}_1^2$$

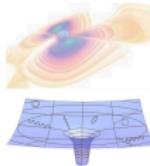
- 3-dim. covariant, as \hat{p}_i is not:

$$\hat{p}_i = p_i - \frac{1}{2} g_{ij} \gamma^{lm} \gamma^{kj}_{,m} \hat{S}_{kl} + \dots$$

- Terms like $Q_{1;k}^{ij} \hat{\delta}_1$ or $Q_1^{ij} \hat{\delta}_{1;k}$ can not appear.
- Kerr metric $\Rightarrow c_1 = -\frac{1}{2}$ and $c_2 = 0$.
- c_3 and c_4 do not contribute to the Hamiltonian.



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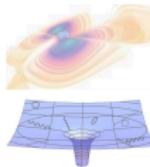


NLO Spin₁-Spin₁ for Black Holes

$$\begin{aligned}H_{S_1^2}^{\text{NLO}} = & \frac{1}{\hat{r}_{12}^3} \left[\frac{m_2}{4m_1^3} (\hat{\mathbf{p}}_1 \cdot \hat{\mathbf{S}}_1)^2 + \frac{3m_2}{8m_1^3} (\hat{\mathbf{p}}_1 \cdot \hat{\mathbf{n}}_{12})^2 \hat{\mathbf{S}}_1^2 - \frac{3m_2}{8m_1^3} \hat{\mathbf{p}}_1^2 (\hat{\mathbf{S}}_1 \cdot \hat{\mathbf{n}}_{12})^2 \right. \\& - \frac{3m_2}{4m_1^3} (\hat{\mathbf{p}}_1 \cdot \hat{\mathbf{n}}_{12}) (\hat{\mathbf{S}}_1 \cdot \hat{\mathbf{n}}_{12}) (\hat{\mathbf{p}}_1 \cdot \hat{\mathbf{S}}_1) - \frac{3}{4m_1 m_2} \hat{\mathbf{p}}_2^2 \hat{\mathbf{S}}_1^2 \\& + \frac{9}{4m_1 m_2} \hat{\mathbf{p}}_2^2 (\hat{\mathbf{S}}_1 \cdot \hat{\mathbf{n}}_{12})^2 + \frac{3}{4m_1^2} (\hat{\mathbf{p}}_1 \cdot \hat{\mathbf{p}}_2) \hat{\mathbf{S}}_1^2 \\& - \frac{9}{4m_1^2} (\hat{\mathbf{p}}_1 \cdot \hat{\mathbf{p}}_2) (\hat{\mathbf{S}}_1 \cdot \hat{\mathbf{n}}_{12})^2 + \frac{3}{4m_1^2} (\hat{\mathbf{p}}_1 \cdot \hat{\mathbf{n}}_{12}) (\hat{\mathbf{p}}_2 \cdot \hat{\mathbf{n}}_{12}) \hat{\mathbf{S}}_1^2 \\& - \frac{3}{2m_1^2} (\hat{\mathbf{p}}_1 \cdot \hat{\mathbf{n}}_{12}) (\hat{\mathbf{p}}_2 \cdot \hat{\mathbf{S}}_1) (\hat{\mathbf{S}}_1 \cdot \hat{\mathbf{n}}_{12}) \\& + \frac{3}{m_1^2} (\hat{\mathbf{p}}_2 \cdot \hat{\mathbf{n}}_{12}) (\hat{\mathbf{p}}_1 \cdot \hat{\mathbf{S}}_1) (\hat{\mathbf{S}}_1 \cdot \hat{\mathbf{n}}_{12}) \\& \left. - \frac{15}{4m_1^2} (\hat{\mathbf{p}}_1 \cdot \hat{\mathbf{n}}_{12}) (\hat{\mathbf{p}}_2 \cdot \hat{\mathbf{n}}_{12}) (\hat{\mathbf{S}}_1 \cdot \hat{\mathbf{n}}_{12})^2 \right] \\& - \frac{m_2}{\hat{r}_{12}^4} \left[\frac{9}{2} (\hat{\mathbf{S}}_1 \cdot \hat{\mathbf{n}}_{12})^2 - \frac{5}{2} \hat{\mathbf{S}}_1^2 + \frac{7m_2}{m_1} (\hat{\mathbf{S}}_1 \cdot \hat{\mathbf{n}}_{12})^2 - \frac{3m_2}{m_1} \hat{\mathbf{S}}_1^2 \right]\end{aligned}$$



since 1558



Thank you for your attention!

